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# Preparation of (100) CeO<sub>2</sub> and (110) YBa<sub>2</sub>Cu<sub>3</sub>O<sub>7- $\delta$ </sub> films by laser chemical vapor deposition

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#### **Abstract**

 $CeO_2$  and  $YBa_2Cu_3O_{7-\delta}$  (YBCO) films were prepared by laser chemical vapor deposition. The (100)-oriented  $CeO_2$  films prepared at 975–1088 K consisted of columnar grains with pyramidal caps of the  $CeO_2$  {111} planes. The deposition rate was 3.0–5.4  $\mu$ m h<sup>-1</sup>. The (110)-oriented YBCO films were prepared on the (100)-oriented  $CeO_2$  films at a deposition temperature of 1004 K. The (110)-oriented YBCO films showed wedge-shaped grains.

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Keywords: A. Film; D. CeO2; Laser CVD; YBCO; Orientation control

### 1. Introduction

The high-temperature superconductor YBa<sub>2</sub>Cu<sub>3</sub>O<sub>7- $\delta$ </sub> (YBCO) has been extensively studied since its discovery in 1987 [1,2]. The electrical transport properties of YBCO are highly anisotropic and depend greatly on the crystallographic orientation [3,4]. Since YBCO can carry a large superconducting current along the CuO<sub>2</sub> plane (*ab*-plane), *c*-axis-oriented YBCO films are required for high-temperature superconducting power cables. YBCO also exhibits a photovoltaic effect [5]. Because the coherence length along the *ab*-plane ( $\xi_{ab}$ =1.2–15 nm) is much larger than that along the *c*-axis ( $\xi_c$ =0.1–0.3 nm) [6,7], (110)-oriented YBCO films are preferable in photovoltaic devices [8].

(110)-oriented YBCO films have been prepared on the (110) planes of single crystal substrates, mainly SrTiO<sub>3</sub> [9–12]. However, (103)-oriented YBCO was often observed with (110)-oriented grains because of the alternating coherent epitaxial relationship between the YBCO (103) plane and substrate. We previously prepared (100)-oriented CeO<sub>2</sub> films on single crystal and polycrystalline substrates by laser chemical vapor deposition (laser CVD) [13,14]. The CeO<sub>2</sub>

films consisted of columnar grains and showed strong (100) orientation. The top of the (100)-oriented  $CeO_2$  columnar grain was terminated by pyramidal caps of the  $CeO_2$  {111} plane. In these studies, a metalorganic (MO) compound of Ce was used as a solid precursor. Although MOCVD process using a solid compound has widely been studied, in terms of preparing multi-component film in industrial applications, a stable and continuous precursor feeding system using a single liquid precursor is feasible for a long-length production of  $CeO_2$  and YBCO films.

In the present study, by utilizing the pyramidal morphology of CeO<sub>2</sub> film as a template, (110)-oriented YBCO films were prepared on (100)-oriented CeO<sub>2</sub> films.

#### 2. Experimental

The laser CVD apparatus used in this study is depicted in Fig. 1. A continuous wave Nd:YAG laser (wavelength: 1064 nm) was employed. CeO<sub>2</sub> films were prepared using Ce(DPM)<sub>4</sub> (dipivaloylmethanate; DPM) dissolved in C<sub>4</sub>H<sub>8</sub>O (tetrahydrofuran; THF) at a concentration of 0.116 mol L<sup>-1</sup>. The source solution was delivered at a supply rate of  $0.25 \times 10^{-4} \, \text{L s}^{-1}$  by a plunger pump. The flow rates of Ar and O<sub>2</sub> gases were 1.52 and 0.34 Pa m<sup>3</sup> s<sup>-1</sup>, respectively. The evaporation chamber was heated to 653 K. A multilayer-coated Hastelloy C276 tape

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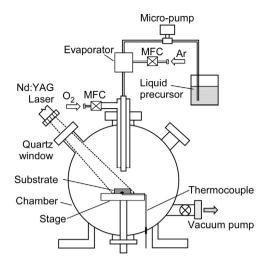


Fig. 1. A schematic of laser CVD apparatus using a single liquid source precursor delivery system.

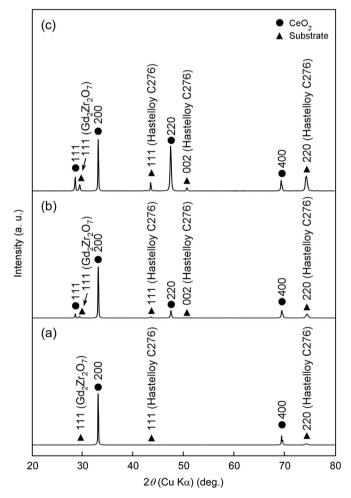


Fig. 2. XRD patterns of  $CeO_2$  films prepared at  $T_{dep} = 1040$  K (a), 1110 K (b) and 1134 K (c).

 $(10 \text{ mm} \times 5 \text{ mm} \times 0.1 \text{ mm})$  with a structure of LaMnO<sub>3</sub>/MgO/Gd<sub>2</sub>Zr<sub>2</sub>O<sub>7</sub>/Hastelloy C276 was used as a substrate and was preheated on a hot stage at 823 K. The total pressure was maintained at 0.8 kPa.

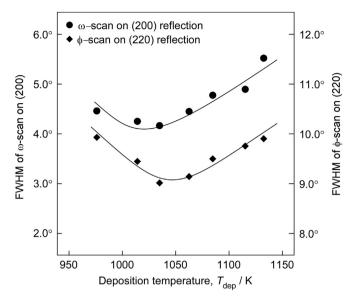


Fig. 3. Effect of  $T_{\rm dep}$  on the FWHM of  $\omega$ -scan on the (200) reflection and that of  $\phi$ -scan on the (220) reflection of the CeO<sub>2</sub> films.

YBCO films were prepared using Y(DPM)<sub>3</sub>, Ba(DPM)<sub>2</sub>/Ba (TMOD)<sub>2</sub> (2,2,6,6-tetramethyl-3,5-octanedionate; TMOD) and Cu(DPM)<sub>2</sub>. These were mixed in a molar ratio of Y:Ba: Cu=1:2.4:2.8 and then dissolved in THF in a concentration of 0.01 mol L<sup>-1</sup> of Y. The source solution was delivered at a supply rate of  $0.11 \times 10^{-4}$  L s<sup>-1</sup>. The substrate preheating temperature was 923 K. The flow rate of O<sub>2</sub> gas was 0.85 Pa m<sup>3</sup> s<sup>-1</sup>.

The phases and orientations of the  $CeO_2$  and YBCO films were identified using an X-ray diffractometer (XRD; Rigaku RAD-2C). The crystallinity of the  $CeO_2$  film was evaluated from the full width at half maximum (FWHM) of the  $\omega$ -scan (rocking curve) of the (200) reflection. The in-plane orientation of the  $CeO_2$  film was evaluated from the FWHM of the  $\phi$ -scan of the (220) reflection, which was measured by a pole figure XRD (Rigaku Ultima IV). The microstructure was observed using a field emission scanning electron microscope (FESEM; JEOL JSM-7500F). A schematic of the crystal structure was drawn using VESTA [15].

#### 3. Results and discussion

The deposition temperature ( $T_{\rm dep}$ ) increased from 975 to 1134 K with laser power ( $P_{\rm L}$ ) from 46 to 132 W. Fig. 2 shows the XRD patterns of the CeO<sub>2</sub> films prepared at  $T_{\rm dep}$ =975–1134 K. (100)-oriented CeO<sub>2</sub> films were obtained at  $T_{\rm dep}$ =975–1088 K (Fig. 2(a)). Small peaks of (111) and (220) orientations were observed in the diffraction patterns of the (100)-oriented CeO<sub>2</sub> films at  $T_{\rm dep}$ =1110–1134 K (Fig. 2(b)).

Fig. 3 shows the effect of  $T_{\rm dep}$  on the FWHM values of the  $\omega$ -scan of the (200) reflection and that of the  $\phi$ -scan of the (220) reflection of the CeO<sub>2</sub> films. With increasing  $T_{\rm dep}$  from 975 to 1134 K, the FWHM values of the  $\omega$ -scan first decreased from 4.3° to 4.1° and then increased to 5.4°, with a minimum of 4.1° at  $T_{\rm dep}$  = 1040 K. The FWHM values of the  $\phi$ -scan of the (220) reflection of the CeO<sub>2</sub> films first decreased from 9.9°

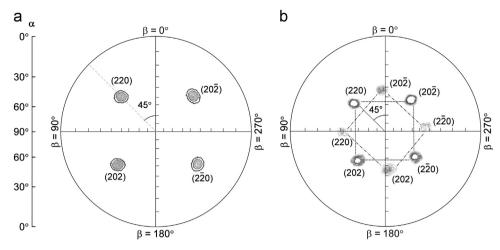


Fig. 4. X-ray pole figure patterns of the (220) reflection of the  $CeO_2$  films prepared at  $T_{dep} = 1040$  K (a) and 1110 K (b).

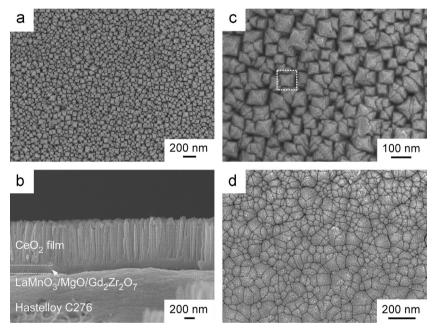


Fig. 5. FESEM images of  $CeO_2$  films prepared at  $T_{dep} = 1040$  K (a)–(c) and 1134 K (d).

to  $9.0^{\circ}$  and then increased to  $9.8^{\circ}$ , with a minimum of  $9.0^{\circ}$  at  $T_{\rm dep}\!=\!1040~\rm K$ . For the  ${\rm CeO_2}$  films prepared on multilayer-coated Hastelloy C273 tape by laser CVD using a solid precursor, the FWHM values of  $\omega$ - and  $\varphi$ -scan were  $4.3^{\circ}$  and  $6.0^{\circ}$ , respectively [14]. Thus, the present  ${\rm CeO_2}$  films prepared using a liquid precursor were comparable to the  ${\rm CeO_2}$  films prepared using a solid precursor.

Fig. 4 shows the XRD pole figures of the (220) reflection of the CeO<sub>2</sub> films prepared at  $T_{\rm dep}{=}1040$  and 1110 K. The pole figure of the (220) reflection of the CeO<sub>2</sub> films prepared at  $T_{\rm dep}{=}975{-}1088$  K showed a fourfold pattern at approximately  $\alpha{=}45^{\circ}$  (Fig. 4(a)). This was attributed to the (220) plane with the complementary angle of  $45^{\circ}$  to the (100) plane. For the CeO<sub>2</sub> films prepared at  $T_{\rm dep}{=}1110$  K (Fig. 4(b)), the pole figures showed two groups of diffraction patterns, which consisted of strong fourfold reflections and weak fourfold reflections. These diffraction patterns were attributed to the

(220) plane with the complementary angle of  $45^{\circ}$  to the (100) plane. The azimuth angle  $\beta$  of  $45^{\circ}$  between strong and weak reflections indicates a  $45^{\circ}$  rotation for a small volume faction of (100)-oriented grains. In the case of the CeO<sub>2</sub> film prepared at  $T_{\rm dep} = 1134$  K, no poles were observed in the diffraction patterns of the (220) reflection because of the dominance of the (111)-oriented grains in the CeO<sub>2</sub> film (Fig. 2(c)).

Fig. 5 shows the microstructures of the  $CeO_2$  films prepared at  $T_{dep} = 1040$  and 1134 K. The low magnification image of the  $CeO_2$  film prepared at  $T_{dep} = 1040$  K shows pyramidal surface grains (Fig. 5(a)). The cross-sectional image shows a columnar growth of the  $CeO_2$  film (Fig. 5(b)). The magnified surface image of the fine grains prepared at  $T_{dep} = 1100$  K (Fig. 5(c)) shows orthogonal pyramidal caps whose four-sided facets were attributed to the {111} planes [16]. Because the {111} planes have the lowest surface energy compared to the (100) and (110) planes [17,18], they would be preferentially formed at

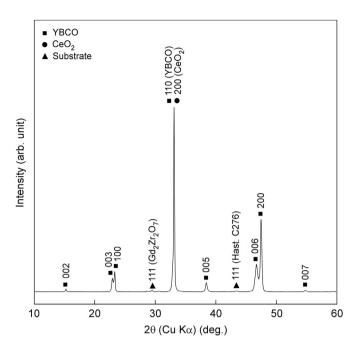


Fig. 6. XRD pattern of the YBCO films prepared at  $T_{\rm dep}\!=\!1016\,\rm K$  on  ${\rm CeO_2}$  film.

the top surface of the  $\text{CeO}_2$  film. The size of columnar grains was approximately 50 nm in width and was finer than the columnar grains obtained in the  $\text{CeO}_2$  films prepared using a solid precursor (approximately 80 nm in width) [14]. The  $\text{CeO}_2$  film prepared at  $T_{\text{dep}} = 1134 \, \text{K}$  consisted of granular grains (Fig. 5(d)). The deposition rate ( $R_{\text{dep}}$ ) of the  $\text{CeO}_2$  film prepared was  $4.8-5.4 \, \mu \text{m h}^{-1}$  at  $T_{\text{dep}} = 975-1040 \, \text{K}$ . A further increase in  $T_{\text{dep}}$  to  $1134 \, \text{K}$  caused a decrease of  $R_{\text{dep}}$  to  $3.6 \, \mu \text{m h}^{-1}$ . The  $R_{\text{dep}}$  of the  $\text{CeO}_2$  films prepared using a liquid precursor was slightly higher than those of films prepared using a solid precursor ( $3.4-4.6 \, \mu \text{m h}^{-1}$ ) [14]. The optimized  $T_{\text{dep}}$  for the (100)-oriented  $\text{CeO}_2$  film was 1040 K (the corresponding  $P_{\text{L}}$  was 77 W), at which the  $\text{CeO}_2$  film showed the highest in-plane orientation (Fig. 3).

The YBCO films were prepared at  $T_{\rm dep} = 1004$  K ( $P_{\rm L} = 85$  W) on the (100)-oriented CeO<sub>2</sub> films. Fig. 6 shows the XRD pattern of the YBCO film, which showed a (110) preferred orientation. The XRD peak of the YBCO (110) plane ( $2\theta = 32.8^{\circ}$ ) overlapped with that of the CeO<sub>2</sub> (200) plane ( $2\theta = 33.1^{\circ}$ ). A small amount of YBCO grains with *a*-axis and *c*-axis orientations was also present in the (110)-oriented YBCO film.

Fig. 7 shows the X-ray pole figure of the (100) reflection of the (110)-oriented YBCO film prepared at  $T_{\rm dep} = 1004~{\rm K}$  on (100)-oriented CeO<sub>2</sub> film. The YBCO (100) reflection showed four spots at approximately  $\alpha = 45^{\circ}$ . This was attributed to the YBCO (100) plane with the complementary angle of  $45^{\circ}$  to the YBCO (110) plane. Because the YBCO (100) plane exhibits twofold reflections for YBCO (110) plane, the fourfold reflections observed in the pole figure suggested that (110)-oriented YBCO grains had a  $90^{\circ}$ -rotated domain (indicated as R in Fig. 7). The in-plane orientation relationship between (110)-oriented YBCO and (100)-oriented CeO<sub>2</sub> films was [001] YBCO || [011] CeO<sub>2</sub>.

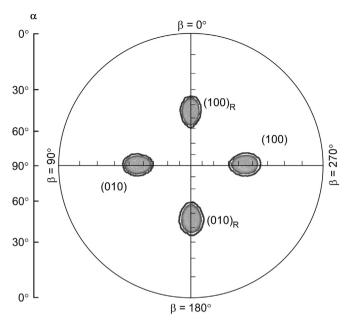


Fig. 7. X-ray pole figure pattern of the YBCO films prepared at  $T_{\rm dep} = 1016$  K on CeO<sub>2</sub> film.

Fig. 8 shows the surface and cross-sectional FESEM images of the YBCO film prepared at  $T_{\rm dep} = 1004$  K on (100)-oriented CeO<sub>2</sub> film. The low magnification surface image of the (110)oriented YBCO film shows a dense surface with crystal grains approximately 200-300 nm in size (Fig. 8(a)). The magnified surface image shows that the (110)-oriented YBCO film consists of crystal grains with wedge-shaped top surfaces. The four facets of these wedge-shaped grains could be the {100} and {001} planes.  $R_{\text{dep}}$  of the (110)-oriented YBCO film was 8.7  $\mu$ m h<sup>-1</sup>. Image (d) depicts the growth relationship between (110)-oriented YBCO grains and the (100)-oriented CeO<sub>2</sub> grains with pyramidal caps of the {111} planes. Between two neighboring CeO<sub>2</sub> grains, a V-shaped valley can be formed along the (110) direction, as shown in Fig. 8(d). At the nucleation stage of the YBCO film, the V-shaped valley could serve as the knit sites for the (110)-oriented YBCO grains to decrease the total surface energy while the YBCO nucleates at these sites [19]. Therefore, (110)-oriented YBCO grains could favorably grow on the (100)-oriented CeO<sub>2</sub> grains. These wedge-shaped grains were perpendicular to each other, as marked in Fig. 8. This grain growth model is coincident with the in-plane orientations derived from the X-ray pole figure analysis (Figs. 4 and 6).

## 4. Conclusions

CeO<sub>2</sub> films with (100) preferred orientation were prepared at a high deposition rate of 3.0–5.4  $\mu$ m h<sup>-1</sup> on multilayer-coated Hastelloy C276 tape by laser CVD. The FWHM of the  $\omega$ -scan of the (200) reflection and that of the  $\phi$ -scan of the (220) reflection were 4.1–5.4° and 9.0–9.9°, respectively. The (100)-oriented CeO<sub>2</sub> films had columnar grains in the cross section. The YBCO film with (110) preferred orientation was prepared on (100)-oriented CeO<sub>2</sub> films at a high

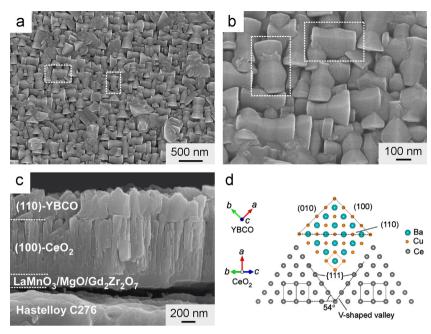


Fig. 8. FESEM surface (a) and cross-sectional (c) images of the (110)-oriented YBCO film prepared at  $T_{\rm dep}$ =1004 K. Image (b) is the magnified surface image. Image (d) depicts the growth relationship between the (110)-oriented YBCO grains and the (100)-oriented CeO<sub>2</sub> grains.

deposition rate of 8.7  $\mu$ m h<sup>-1</sup>. The (110)-oriented YBCO film showed wedge-shaped grains.

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